

ZCE 111  
Assignment 12

# Q1: SK2 code for forced pendulum

Develop a code to implement SK2 for the case of a forced pendulum experiencing no drag force but a driving

force  $F_D \sin(\Omega_D t)$ ,  $\Omega = \sqrt{g/l}$ ,  $l = 1.0$  m,  $m = 1$  kg;  $F_D = 1$  N;  $\Omega_D = 0.99 \Omega$ ;

Boundary conditions:  $\theta(0) = 0.0$ ;  $\omega(t = 0) = 0$ ;

$$\frac{d^2 \theta}{dt^2} = -\frac{g}{l} \theta - q \frac{d\theta}{dt} + \frac{F_D \sin(\Omega_D t)}{m l}$$

## Q2: Stability of the total energy a SHO in RK2.

$\omega = \frac{d\theta}{dt}$ , angular velocity.  $m=1\text{kg}$ ;  $l=1\text{m}$ .

The total energy of the SHO in can be calculated as

$$\begin{aligned} E_{i+1} &= K_{i+1} + U_{i+1} = \frac{1}{2}m(l\omega_{i+1})^2 + mgl(1 - \cos\theta_{i+1}) \\ &\approx \frac{1}{2}ml^2\omega_{i+1}^2 + mgl\left[1 - \left(1 - \frac{\theta_{i+1}^2}{2}\right)\right] \\ &= \frac{1}{2}ml^2\omega_{i+1}^2 + \frac{1}{2}mgl\theta_{i+1}^2 \end{aligned}$$

User your RK2 code to track the total energy for  $t$  running from  $t=0$  till  $t=25T$ ;  $T=\sqrt{g/l}$ . Boundary conditions:  $\omega(0) = \sqrt{\frac{g}{l}}$ ;  $\theta(0) = 0$ . Check that indeed  $E_i$  should remain constant throughout all  $t_i$ .

### Q3: Planetary motion in polar coordinates

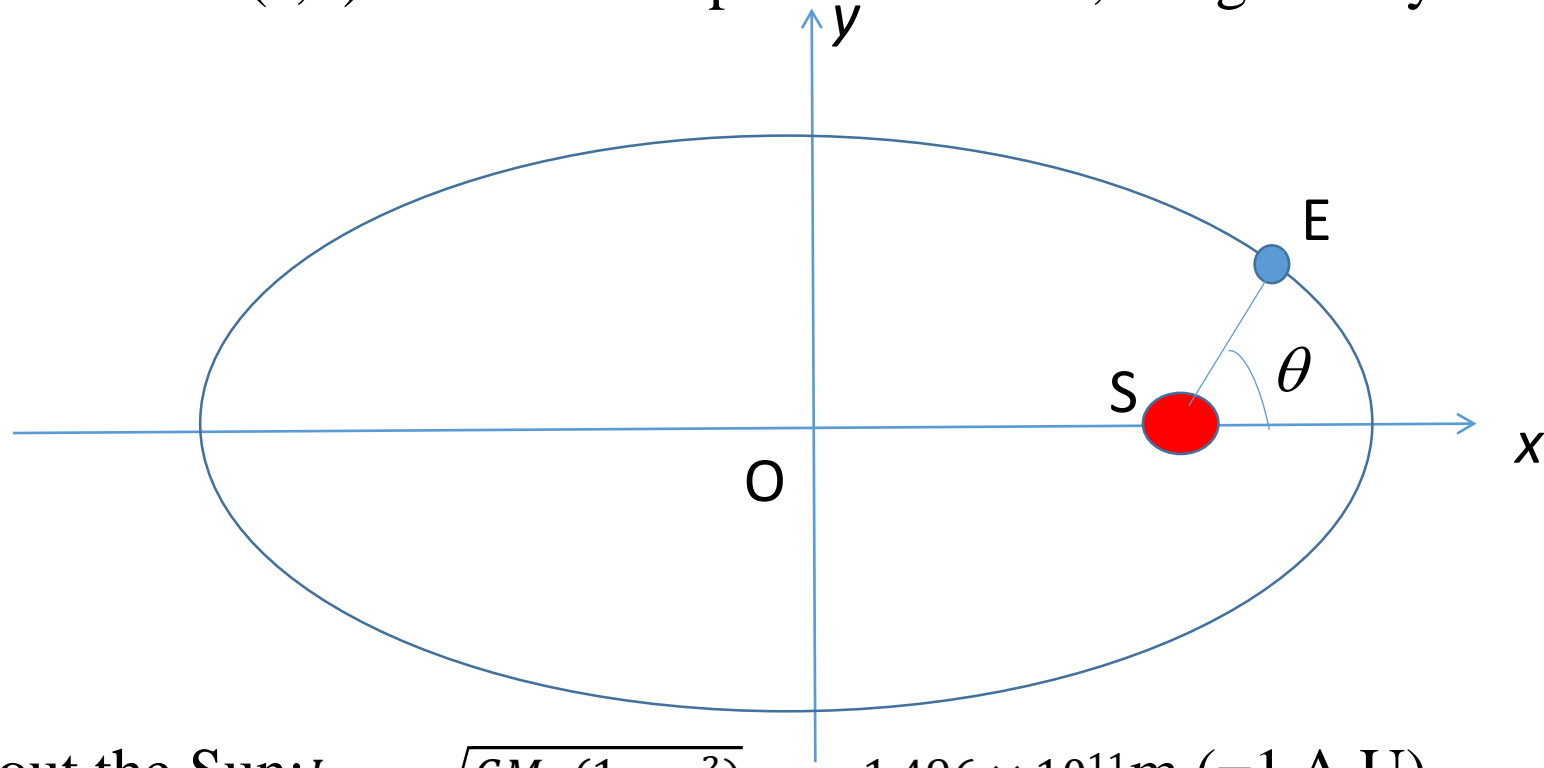
The Earth (E) moves in an elliptical orbit around the Sun (S). In polar coordinate, the equation of motions for the radial and angular coordinate  $(r, \theta)$  of E with respect to the Sun, are given by

$$m\ddot{r} - mr\left(\frac{L}{mr^2}\right)^2 = -\frac{GMm}{r^2}$$

$$mr^2\dot{\theta} = L$$

where

$$\ddot{r} = \frac{d^2r}{dt^2}; \dot{\theta} = \frac{d\theta}{dt}$$



$L$  is the angular momentum of Earth about the Sun:  $L = m\sqrt{GMa(1 - \epsilon^2)}$ ;  $a = 1.496 \times 10^{11}\text{m}$  (=1 A.U) (semimajor);  $m = 5.97 \times 10^{24}\text{kg}$ , mass of Earth;  $M = 1.987 \times 10^{30}\text{kg}$ , Mass of the Sun;

$G = 6.673 \times 10^{-11}\text{Nm}^2/\text{kg}^2$ ;  $\epsilon = 0.017$  (eccentricity). The period is  $T = 2\pi\sqrt{\frac{a^3}{GM}}$

### Q3: Planetary motion in polar coordinates (cont.)

- (i) Solve the radial equation,  $m\ddot{r} - mr \left(\frac{L}{mr^2}\right)^2 = -\frac{GMm}{r^2}$ , numerically using your RK2 code, so that you can plot the graph of  $r$  as a function of time for  $t$  from 0 to the period of the orbit,  $T$  ( $= 1$  year). Boundary conditions:  $r(0) = \frac{L^2}{GMm^2} \frac{1}{(1+\epsilon)}$ ,  $\dot{r}(0) = 0$ .
- (ii) Hence, plot the angular velocity function  $\dot{\theta}(t)$  as the time varies throughout the year.
- (iii) Plot the normalised angular function  $\frac{\theta(t)}{2\pi}$  as the time varies throughout the whole period.  
(\*Think carefully how you are going to do this using the prior knowledge you have acquired. Think Euler.)
- (iv) Visualise the evolution of the Earth's location,  $(x(t), y(t))$  on the orbit on a  $x$ - $y$  plot using the Manipulate function. Your visualization should contain the origin,  $x$ - and  $y$ -axes.

## Q4: Planetary motion in polar coordinates, for Halle's comet

Repeat Q2 for Halle's comet. The parameters of Halle's comet are:

$$a = 17.8 \text{ A. U}; \quad m = 2.2 \times 10^{14} \text{ kg}; \quad \epsilon = 0.967.$$

Check that the period is indeed  $\approx 75$  years.

## Q5: Coulomb charge scattering in 3D

Develop a code based on velocity Verlet algorithm (see the sample code [verlet\\_algorithm\\_2D\\_coulomb\\_scatterings.nb](#) as a reference) to simulate the scattering process of a fixed charge  $Q = +1$  unit, located at  $(r_{20}\sin\theta_2\sin\phi_2, r_{20}\cos\phi_2, r_{20}\sin\theta_2\cos\phi_2)$ , by a moving charge  $q = +1$  unit (with mass  $m = 1$  unit), initially located at  $(0,0,0)$ , and is shooting towards  $Q$  with an initial speed  $v_0 = 1$  unit. In your simulation, set  $k = \frac{1}{4\pi\epsilon_0} = 1$ ,  $r_{20} = 1$  unit,  $\theta_2 = \phi_2 = 45^\circ$ .

## Q6: Charge moving in a magnetic field

A charge (mass  $m$  and charge  $q$ ) moving with velocity  $\mathbf{v} = (v_x, v_y, v_z)$  in a magnetic field  $\mathbf{B} = (B_x, B_y, B_z)$  experiences a velocity-dependent Lorentz force  $\mathbf{F} = (F_x, F_y, F_z) = q \mathbf{v} \times \mathbf{B}$ . Develop a code based on the Störmer-Verlet integration algorithm to simulate the dynamical path of the charge particle moving through the magnetic field.

Assume:  $q = +1$  unit, mass  $m = 1$  unit, initially located at  $(0, 0, 0)$ , initial velocity  $(v_{0x}, v_{0y}, v_{0z})$ ,  $v_{0x} = v_{0y} = 0.1$  unit,  $v_{0z} = 0.05$  unit,  $\mathbf{B} = (0, 0, B_z)$ ,  $B_z = 0.1$  unit. You should see a helical trajectory circulating about the  $z$ -direction.

(my code: `verlet_algorithm_3D_coulomb_helix.nb`)