ZCT 104 (Class A)

MODERN PHYSICS

ACADEMIC SESSION 2008/09 (SECOND SEMESTER)

LECTURE NOTES, TUOTRIAL PROBLEM SET And PAST YEAR QUESTIONS

> By Yoon Tiem Leong School Of Physics USM, Penang

POWERPOINT LECTURE NOTES SESSI 2008/09 SEMESTER II

Special theory of Relativity

<u>Notes based on</u> Understanding Physics by Karen Cummings et al., John Wiley & Sons

















In a "covered" reference frame, we can't tell whether we are moving or at rest

• Without referring to an external reference object (such as a STOP sign or a lamp post), whatever experiments we conduct in a constantly moving frame of reference (such as a car at rest or a car at constant speed) could not tell us the state of our motion (whether the reference frame is at rest or is moving at constant velocity)



Physical laws must be invariant in any reference frame

- Such an inability to deduce the state of motion is a consequence of a more general principle:
- There must be no any difference in the physical laws in any reference frame with constant velocity
- (which would otherwise enable one to differentiate the state of motion from experiment conducted in these reference frame)
- Note that a reference frame at rest is a special case of reference frame moving at constant velocity (v = 0 = constant)



Einstein's Puzzler about running fast while holding a mirror



- Says Principle of Relativity: Each fundamental constants must have the same numerical value when measured in any reference frame (c, h, e, m_e, etc)
- (Otherwise the laws of physics would predict inconsistent experimental results in different frame of reference which must not be according to the Principle)
- Light always moves past you with the same speed *c*, no matter how fast you run
- Hence: you will not observe light waves to slow down as you move faster 13







Touchstone Example 38-1: Communication storm!

• A sunspot emits a tremendous burst of particles that travels toward the Earth. An astronomer on the Earth sees the emission through a solar telescope and issues a warning. The astronomer knows that when the particle pulse arrives it will wreak havoc with broadcast radio transmission. Communications systems require ten minutes to switch from over-the-air broadcast to underground cable transmission. What is the maximum speed of the particle pulse emitted by the Sun such that the switch can occur in time, between warning and arrival of the pulse? Take the sun to be 500 light-seconds distant from the Earth.

Solution		
•	It takes 500 seconds for the warning light flash to travel the distance of 500 light-seconds between the Sun and the Earth and enter the astronomer's telescope. If the particle pulse moves at half the speed of light, it will take twice as long as light to reach the Earth. If the pulse moves at one-quarter the speed of light, it will take four times as long to make the trip. We generalize this by saying that if the pulse moves with speed v/c , it will take time to make the trip given by the expression:	
•	$\Delta t_{\text{pulse}} = 500 \text{ s/} (v_{\text{pulse}}/c)$ How long a warning time does the Earth astronomer have between arrival of the light flash carrying information about the pulse the arrival of the pulse itself? It takes 500 seconds for the light to arrive. Therefore the warning time is the difference between pulse transit time and the transit time of light:	
•	$\Delta t_{\text{warning}} = \Delta t_{\text{pulse}} - 500 \text{ s.}$ But we know that the minimum possible warning time is 10 min = 600 s.	
•	Therefore we have $(0) = 500 = 100 = 5000 = 500 = 500 = 500 = 500 = 500 = 500 = 500 = 500$	
•	which gives the maximum value for v_{puls} if there is to he sufficient time for warning:	
	$v_{\text{puls}} = 0.455 \ c.$ (Answer)	
•	Observation reveals that pulses of particles emitted from the sun travel much slower than this maximum value. So we would have much longer warning time than calculated here.	





















Example 38-2: Simultaneity of the Two Towers Frodo is an intelligent observer • standing next to Tower A, which emits a flash of light every 10 s. 100 km distant from him is the tower B, stationary with respect to him, that also emits a light flash every 10 s. Frodo wants to know whether or not each flash is emitted from remote tower B simultaneous with (at the same time as) the flash from Frodo's own Tower A. Explain how to do this with out leaving Frodo position next to Tower A. Be specific and use numerical values. 29

	Solution	
•	Frodo is an intelligent observer, which means that he kno how to take into account the speed of light in determining the time of a remote event, in this case the time of emissio of a flash by the distant Tower B. He measures the time lapse between emission of a flash by his Tower A and his reception of flash from Tower B.	w on
•	If this time lapse is just that required for light move from Tower B to Tower A, then the two emissions occur the same time.	
•	The two Towers are 100 km apart. Call this distance L . Then the time t for a light flash to move from B to A is	
•	$t = L/c = 10^5 \text{ m/3} \times 10^8 \text{ m/s} = 0.333 \text{ ms.}$ (ANS)	
•	If this is the time Frodo records between the flash nearby Tower A and reception of the flash from distant tower then he is justified in saying that the two Towers emit flashes simultaneously in his frame.	n 30

One same event can be considered in any frame of reference

- One same event, in principle, can be measured by many separate observers in different (inertial) frames of reference (reference frames that are moving at a constant velocity with respect to each other)
- Example: On the table of a moving train, a cracked pot is dripping water
- The rate of the dripping water can be measured by (1) Ali, who is also in the train, or by (2) Baba who is an observer standing on the ground. Furthermore, you too can imagine (3) ET is also performing the same measurement on the dripping water from Planet Mars. (4) By Darth Veda from Dead Star.³¹















Time dilation as direct consequence of constancy of light speed

- According to the Principle of Relativity, the speed of light is invariant (i.e. it has the same value) in every reference frame (constancy of light speed)
- A direct consequence of the constancy of the speed of light is time stretching
- Also called time dilation
- Time between two events can have different values as measured in lab frame and rocket frames in relative motion

39

"Moving clock runs slow"



RE 38-5

• Suppose that a beam of pions moves so fast that at 25 meters from the target in the laboratory frame exactly half of the original number remain undecayed. As an experimenter, you want to put more distance between the target and your detectors. You are satisfied to have one-eighth of the initial number of pions remaining when they reach your detectors. How far can you place your detectors from the target?

41

• ANS: 75 m











RE 38-6 • A set of clocks is assembled in a stationary boxcar. They include a quartz wristwatch, a balance wheel alarm clock, a pendulum grandfather clock, a cesium atomic clock, fruit fies with average individual lifetimes of 2.3 days. A clock based on radioactive decay of nuclei, and a clock timed by marbles rolling down a track. The clocks are adjusted to in at the same rate as one another. The boxcar is then gently accelerated along a smooth horizontal track to a final velocity of 300 km/hr. At this constant final speed, which clocks will run at a different rate from the others as measured in that moving boxcar.



"the invariant space-time interval"

- We call the RHS, $s^2 \equiv (c\Delta t)^2 (\Delta x)^2$ "invariant space-time interval squared" (or sometimes simply "the space-time interval")
- In words, the space-time interval reads:
- $s^2 = (c \times \text{time interval between two events as observed in the frame})^2$ -(distance interval between the two events as observed in the frame)²
- We can always calculate the space-time intervals for any pairs of events
- The interval squared s^2 is said to be an *invariant* because it has the same value as calculated by all observers (take the simile of the mass-to-high² ratio)
- Obviously, in the light-clock gadanken experiment, the space-time interval of the two light pulse events $s^2 \equiv (c\Delta t)^2 (\Delta x)^2 = (c\Delta \tau)^2$ is positive because $(c\Delta \tau)^2 > 0$
- The space-time interval for such two events being positive is deeply related to the fact that such pair of events are causally related







Proper time

- Imagine you are in the rocket frame, O', observing two events taking place at the same spot, separated by a time interval Δτ (such as the emission of the light pulse from source (EV1), and re-absorption of it by the source again, (EV2))
- Since both events are measured on the same spot, they appeared at rest wrp to you
- The time lapse Δτ between the events measured on the clock at rest is called the proper time or wristwatch time (one's own time)






























Why?

• This is due to the invariance of the space-time invariant in all frames, (i.e. the invariant must have the same value for all frames)

How invariance of space-time interval explains disagreement on simultaneity by two observers

- Consider a pair of events with space-time interval $s^2 = (c\Delta t)^2 (\Delta x)^2 = (c\Delta t')^2 (\Delta x')^2$
- where the primed and un-primed notation refer to space and time coordinates of two frames at relative motion (lets call them O and O')
- Say O observes two simultaneous event in his frame (i.e. $\Delta t = 0$) and are separate by a distance of (Δx) , hence the space-time interval is $s^2 = -(\Delta x)^2$
- The space-time interval for the same two events observed in another frame, O', $s'^2 = (c\Delta t')^2 - (\Delta x')^2$ must read the same, as - $(\Delta x)^2$
- Hence, $(c\Delta t')^2 = (\Delta x')^2 (\Delta x)^2$ which may not be zero on the RHS. i.e. $\Delta t'$ is generally not zero. This means in frame O', these events are not observed to be occurring simultaneously 70





Solution

- Music has been emitted from the tape player. This is a fact that must be true in both frames of reference. Hence Sam on the ground will be able to hear the music (albeit with some distortion).
- When the meet for coffee, they will both agree that some tape has been wound from one spool to the other in the tape recorder.

Touchstone Example 38-5: Principle of Relativity Applied Divide the following items into two lists, On one list, labeled SAME, place

- Divide the following items into two lists, On one list, labeled SAME, place items that name properties and laws that are always the same in every frame. On the second list, labeled MAY BE DIF FERENT. place items that name properties that can be different in different frames:
- a. the time between two given events
- b. the distance between two given events
- c. the numerical value of Planck's constant h
- d. the numerical value of the speed of light c
- e. the numerical value of the charge e on the electron
- f. the mass of an electron (measured at rest)
- g. the elapsed time on the wristwatch of a person moving between two given events
- h. the order ot elements in the periodic table
- i. Newton's First Law of Motion ("A particle initially at rest remains at rest, and ...")
- j. Maxwell's equations that describe electromagnetic fields in a vacuum
- k. the distance between two simultaneous events

74

























In classical mechanics, mechanical energy (kinetic + potential) of an object is closely related to its momentum and mass

Since in SR we have redefined the classical mass and momentum to that of relativistic version

$$m_{\text{class}}(\text{cosnt}, = m_0) \rightarrow m_{\text{SR}} = m_0 \gamma$$

$$p_{\text{class}} = m_{\text{class}} \nu \rightarrow p_{\text{SR}} = (m_0 \gamma) \nu$$

we must also modify the relation btw work and energy so that the law conservation of energy is consistent with SR

•E.g, in classical mechanics, $K_{class} = p^2/2m = mv^2/2$. However, this relationship has to be supplanted by the relativistic version

87

$$K_{class} = mv^2/2 \rightarrow K_{SR} = E - m_0 c^2 = \gamma m_0 c^2 - m_0 c^2$$

 \otimes We shall derive *K* in SR in the following slides



Derivation of relativistic kinetic
energy
Force = rate change of
momentum

$$W = \int_{x_1=0}^{x_2} F dx = \int_{x_1=0}^{x_2} \frac{dp}{dt} dx = \int_{x_1=0}^{x_2} \left(\frac{dp}{dx} \frac{dx}{dt}\right) dx$$

 $= \int_{x_1=0}^{x_2} \frac{dp}{dx} v dx = \int_{x_1=0}^{x_2} \left(\frac{dp}{dv} \frac{dv}{dx}\right) v dx = \int_{0}^{v} \frac{dp}{dv} v dv$
where, by definition, $v = \frac{dx}{dt}$ is the velocity of the
object

Explicitly,
$$p = \gamma m_0 v$$
,
Hence, $dp/dv = d/dv (\gamma m_0 v)$
 $= m_0 [v (d\gamma/dv) + \gamma]$
 $= m_0 [\gamma + (v^2/c^2) \gamma^3] = m_0 (1 - v^2/c^2)^{-3/2}$
in which we have inserted the relation
 $\frac{d\gamma}{dv} = \frac{d}{dv} \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{v}{c^2} \frac{1}{\left(1 - \frac{v^2}{c^2}\right)^{3/2}} = \frac{v}{c^2} \gamma^3$
 $W = m_0 \int_0^v v \left(1 - \frac{v^2}{c^2}\right)^{-3/2} dv$
integrate
 $M = m_0 \gamma c^2 - m_0 c^2 = mc^2 - m_0 c^2$











Example 38-6: Energy of Fast Particle

• A particle of rest mass m_0 moves so fast that its total (relativistic) energy is equal to 1.1 times its rest energy.

- (a) What is the speed v of the particle?
- (b) What is the kinetic energy of the particle?























Solution

- (i) $K = 2mc^2 2m_0c^2 = 2(\gamma 1)m_0c^2$
- (ii) $E_{\text{before}} = E_{\text{after}} \Rightarrow 2\gamma m_0 c^2 = Mc^2 \Rightarrow M = 2\gamma m_0$
- Mass increase $\Delta M = M 2m_0 = 2(\gamma 1)m_0$
- Approximation: $v/c = ...=1.5 \times 10^{-6} \Rightarrow \gamma \approx 1 + \frac{1}{2} \frac{v^2}{c^2}$ (binomail expansion) $\Rightarrow M \approx 2(1 + \frac{1}{2} \frac{v^2}{c^2})m_0$
- Mass increase $\Delta M = M 2m_0$

$$\approx (v^2/c^2)m_0 = (1.5 \times 10^{-6})^2 m_0$$

- Comparing *K* with ΔMc^2 : the kinetic energy is not lost in relativistic inelastic collision but is converted into the mass of the final composite object, i.e. kinetic energy is conserved
- In contrast, in classical mechanics, momentum is conserved but kinetic energy is not in an inelastic collision

Relativistic momentum and
relativistic Energy
In terms of relativistic momentum, the relativistic
total energy can be expressed as followed
$$E^{2} = \gamma^{2}m_{0}^{2}c^{4}; p^{2} = \gamma^{2}m_{0}^{2}v^{2} \Rightarrow \frac{v^{2}}{c^{2}} = \frac{c^{2}p^{2}}{E^{2}}$$
$$\Rightarrow E^{2} = \gamma^{2}m_{0}^{2}c^{4} = \frac{m_{0}^{2}c^{4}}{1 - \frac{v^{2}}{c^{2}}} = \left(\frac{m_{0}^{2}c^{4}E^{2}}{E^{2} - c^{2}p^{2}}\right)$$
$$\boxed{E^{2} = p^{2}c^{2} + m_{0}^{2}c^{4}} \qquad \text{Energy-momentum}$$







$$Plug p_{\mu}^{2}c^{2} = (K_{\mu} + m_{\mu}c^{2})^{2} - m_{\mu}^{2}c^{4} \text{ into}$$

$$m_{\mu}c^{2} = \sqrt{m_{\mu}^{2}c^{4} + c^{2}p_{\mu}^{2}} + cp_{\mu}$$

$$= \sqrt{m_{\mu}^{2}c^{4} + [(K_{\mu} + m_{\mu}c^{2})^{2} - m_{\mu}^{2}c^{4}]} + \sqrt{(K_{\mu} + m_{\mu}c^{2})^{2} - m_{\mu}^{2}c^{4}}$$

$$= (K_{\mu} + m_{\mu}c^{2}) + \sqrt{(K_{\mu}^{2} + 2K_{\mu}m_{\mu}c^{2})}$$

$$= (4.6MeV + \frac{106MeV}{c^{2}}c^{2}) + \sqrt{(4.6MeV)^{2} + 2(4.6MeV)(\frac{106MeV}{c^{2}})c^{2}}$$

$$= 111MeV + \sqrt{996}MeV = 143MeV$$





Different frame uses different notation for coordinates

O' frame uses {x',y',z';t'} to denote the coordinates of an event, whereas O frame uses {x,y,z;t}

- How to related $\{x',y',z',t'\}$ to $\{x,y,z;t\}$?
- In Newtonian mechanics, we use Galilean transformation



Galilean transformation (applicable only for v << c)

- For example, the spatial coordinate of the object M as observed in O is x and is being observed at a time t, whereas according to O', the coordinate for the space and time coordinates are x' and t'. At low speed v <<c, the transformation that relates {x',t'} to {x,t} is given by Galilean transformation
- {x'=x-vt, t' = t} (x' and t' in terms of x,t)
- $\{x = x' + vt, t = t'\}$ (x and t in terms of x',t')



However, GT contradicts the SR postulate when *v* approaches the speed of light, hence it has to be supplanted by a relativistic version of transformation law when near-to-light speeds are involved: Lorentz transformation (The contradiction becomes more obvious if GT on velocities, rather than on space and time, is considered)



























Simply permute the role of x and x'
and reverse the sign of v
$$t \leftrightarrow t', x \leftrightarrow x', v \rightarrow -v$$
$$x' = \gamma(x - vt) \rightarrow x = \gamma(x' + vt')$$
$$t' = \gamma \left[t - (v/c^2)x \right] \rightarrow t = \gamma \left[t' + (v/c^2)x' \right]$$
The two transformations above are equivalent; use which is
appropriate in a given question







<section-header><text>




















Comparing the LT of velocity with that of GT

Lorentz transformation of velocity:

$$u'_x = \frac{dx'}{dt'} = \frac{u_x - v}{1 - \frac{u_x v}{2}}$$

Galilean transformation of velocity:

$$u'_x = u_x - v$$

LT reduces to GT in the limit $u_x v \ll c^2$





How to express
$$u_x$$
 in terms of u_x' ?• Simply permute v with $-v$ and change the observed u_x with that of u_x' : $u_x \rightarrow u_x, u_x' \rightarrow u_x, v \rightarrow -v$ $u_x = \frac{u_x - v}{1 - \frac{u_x v}{2}} \rightarrow u_x = \frac{u_x' + v}{1 + \frac{u_x' v}{2}}$

Recap: Lorentz transformation
relates

$$\{x',t'\} \nleftrightarrow \{x,t\}; u'_x \nleftrightarrow u_x$$

 $x' = \gamma(x-vt) \quad t' = \gamma [t - (v/c^2)x]$
 $u'_x = \frac{u_x - v}{1 - \frac{u_x v}{c^2}}$
 $x = \gamma(x'+vt') \quad t = \gamma [t'+(v/c^2)x']$
 $u_x = \frac{u'_x + v}{1 + \frac{u'_x v}{c^2}}$

RE 38-12

• A rocket moves with speed 0.9*c* in our lab frame. A flash of light is sent toward from the front end of the rocket. Is the speed of that flash equal to 1.9 *c* as measured in our lab frame? If not, what is the speed of the light flash in our frame? Verify your answer using LT of velocity formula.

Example (relativistic velocity addition)

 Rocket 1 is approaching rocket 2 on a head-on collision course. Each is moving at velocity 4c/5 relative to an independent observer midway between the two. With what velocity does rocket 2 approaches rocket 1?

156



- Choose the observer in the middle as in the stationary frame, O
- Choose rocket 1 as the moving frame O'
- Call the velocity of rocket 2 as seen from rocket 1 u'x. This is the quantity we are interested in
- Frame O' is moving in the +ve direction as seen in O, so v = +4c/5
- The velocity of rocket 2 as seen from O is in the
- -ve direction, so ux = -4c/5
- Now, what is the velocity of rocket 2 as seen from frame O', u'x = ? (intuitively, u'x must be in the negative direction)

See O'approaching from left, hence I see the Velocit : + 4 6 ; v Sh on right, TZF hence UX=-== rocket 2 cket 1 >tre amo Ve 1









PROPERTIES OF WAVES AND MATTERBLACK BODY RADIATION



Interactions

- Matter and energy exist in various forms, but they constantly transform from one to another according to the law of physics
- we call the process of transformation from one form of energy/matter to another energy/matter as 'interactions'
- Physics attempts to elucidate the interactions between them
- But before we can study the basic physics of the matter-energy interactions, we must first have some general idea to differentiate between the two different modes of physical existence: matter and wave

3

• This is the main purpose of this lecture







What is not a `particle'?

• Waves - electromagnetic radiation (light is a form of electromagnetic radiation), mechanical waves and matter waves is classically thought to not have attributes of particles as mentioned

Analogy

- Imagine energy is like water
- A cup containing water is like a particle that carries some energy within it
- Water is contained within the cup as in energy is contained in a particle.
- The water is not to be found outside the cup because they are all retained inside it. Energy of a particle is corpuscular in the similar sense that they are all inside the carrier which size is a finite volume.
- In contrast, water that is not contained by any container will spill all over the place (such as water in the great ocean). This is the case of the energy carried by wave where energy is not concentrated within a finite volume but is spread throughout the space





- For the case of a particle we can locate its location and momentum precisely
- But how do we 'locate' a wave? ٠
- Wave spreads out in a region of space and is not located in any specific point in space like the case of a particle
- To be more precise we says that a plain wave exists within ٠ some region in space, Δx
- For a particle, Δx is just the 'size' of its dimension, e.g. Δx for an apple is 5 cm, located exactly in the middle of a square table, x = 0.5 m from the edges. In principle, we can determine the position of x to infinity

11

But for a wave, Δx could be infinity

In fact, for the `pure' (or 'plain') wave which has 'sharp' wavelength and frequency mentioned in previous slide, the Δx is infinity



A pure wave has $\Delta x \rightarrow$ infinity

- If we know the wavelength and frequency of a pure wave with infinite precision (= the statement that the wave number and frequency are 'sharp'), one can shows that :
- The wave cannot be confined to any restricted region of space but must have an infinite extension along the direction in which it is propagates
- In other words, the wave is 'everywhere' when its wavelength is 'sharp'

13

• This is what it means by the mathematical statement that " Δx is infinity"









• can also be expressed in an equivalence form $\Delta t \Delta \nu \geq 1$

via the relationship $c = v\lambda$ and $\Delta x = c\Delta t$

- Where Δt is the time required to measure the frequency of the wave
- The more we know about the value of the frequency of the wave, the longer the time taken to measure it
- If u want to know exactly the precise value of the frequency, the required time is $\Delta t = infinity$
- We will encounter more of this when we study the Heisenberg uncertainty relation in quantum physics



- We have already shown that the 1-D plain wave is infinite in extent and can't be properly localised (because for this wave, $\Delta x \rightarrow infinity$)
- However, we can construct a relatively localised wave (i.e., with smaller Δx) by :
- adding up two plain waves of slightly different wavelengths (or equivalently, frequencies)















Why are waves and particles so important in physics?

- Waves and particles are important in physics because they represent the only modes of energy transport (interaction) between two points.
- E.g we signal another person with a thrown rock (a particle), a shout (sound waves), a gesture (light waves), a telephone call (electric waves in conductors), or a radio message (electromagnetic waves in space).



(ii) waves and particle, in which a particle gives up all or part of its energy to generate a wave, or when all or part of the energy carried by a wave is absorbed/dissipated by a nearby particle (e.g. a wood chip dropped into water, or an electric charge under acceleration, generates EM wave)



This is an example where particle is interacting with wave; energy transform from the electron's K.E. to the energy propagating in the form of EM wave wave

<section-header><list-item><list-item><list-item>

























Attempt to understand the origin of radiation from hot bodies from classical theories

- In the early years, around 1888 1900, light is understood to be EM radiation
- Since hot body radiate EM radiation, hence physicists at that time naturally attempted to understand the origin of hot body in terms of classical EM theory and thermodynamics (which has been well established at that time)







Total radiated power per unit area

• The total power radiated per unit area (intensity) of the BB is given by the integral

$$I(T) = \int_{0}^{\infty} R(\lambda, T) d\lambda$$

• For a blackbody with a total area of *A*, its total power emitted at temperature *T* is

$$P(T) = AI(T)$$

• Note: The SI unit for *P* is Watt, SI unit for *I* is Watt per meter square; for *A*, the SI unit is meter square

Introducing idealised black body In reality the spectral distribution of intensity of radiation of a given body could depend on the type of the surface which may differ in absorption and radiation efficiency (i.e. frequency-dependent) This renders the study of the origin of radiation by hot bodies case-dependent (which means no good because the conclusions made based on one body cannot be applicable to other bodies that have different surface absorption characteristics) E.g. At the same temperature, the spectral distribution by the exhaust pipe from a Proton GEN2 and a Toyota Altis is different




- Any radiation striking the HOLE enters the cavity, trapped by reflection until is absorbed by the inner walls
- The walls are constantly absorbing and emitting energy at thermal EB
- The nature of the radiation leaving the cavity through the hole depends only on the temperature of the cavity and not the detail of the surfaces nor frequency of the radiation























The average energy per standing wave, $\langle \varepsilon \rangle$

- Theorem of equipartition of energy (a mainstay theorem from statistical mechanics) says that the average energy per standing wave is
- $\langle \varepsilon \rangle = kT$
 - $k = 1.38 \times 10^{-23}$ J/K, Boltzmann constant
- In classical physics, (ɛ) can take any value CONTINOUSLY and there is not reason to limit it to take only discrete values
- (this is because the temperature *T* is continuous and not discrete, hence ε must also be continuous)











Planck's Theory of Blackbody Radiation

- In 1900 Planck developed a theory of blackbody radiation that leads to an equation for the intensity of the radiation
- This equation is in complete agreement with experimental observations

Planck's Wavelength Distribution Function

• Planck generated a theoretical expression for the wavelength distribution (radiance) $R(\lambda T) = \frac{2\pi hc^2}{2\pi hc^2}$

$$\mathcal{K}(\Lambda, \Gamma) = \frac{1}{\lambda^5 \left(e^{hc/\lambda kT} - 1 \right)}$$

- $h = 6.626 \text{ x } 10^{-34} \text{ J} \cdot \text{s}$
- *h* is a fundamental constant of nature

















Example: quantised oscillator vs classical oscillator

- A 2.0 kg block is attached to a massless spring that has a force constant *k*=25 N/m. The spring is stretched 0.40 m from its EB position and released.
- (A) Find the total energy of the system and the frequency of oscillation according to classical mechanics.

79

Solution • In classical mechanics, $E = \frac{1}{2}kA^2 = ... 2.0 \text{ J}$ • The frequency of oscillation is $f = \frac{1}{2\pi}\sqrt{\frac{k}{m}} = ... = 0.56 \text{ Hz}$













CMBR – the most perfect Black Body

- Measurements of the cosmic microwave background radiation allow us to determine the temperature of the universe today.
- The brightness of the relic radiation is measured as a function of the radio frequency. To an excellent approximation it is described by a thermal of blackbody distribution with a temperature of T=2.735 degrees above absolute zero.
- This is a dramatic and direct confirmation of one of the predictions of the Hot Big Bang model.
- The COBE satellite measured the spectrum of the cosmic microwave background in 1990, showing remarkable agreement between theory and experiment.



<section-header><list-item><list-item><list-item><list-item><list-item>













$$(\text{continue}) \text{ how does } E_n = nhf, \\ n=0,1,2,3,\dots \text{ leads to } \langle \varepsilon \rangle = \frac{hv}{e^{hv/kT}-1}?$$
$$= \int_{0}^{\infty} N(n)E_n = \int_{0}^{\infty} N_0 nhv \exp\left(-\frac{nhv}{kT}\right) \\ \int_{0}^{\infty} N(n) = \int_{0}^{\infty} N_0 \exp\left(-\frac{nhv}{kT}\right) \\ = \int_{0}^{\infty} N(n) = \int_{0}^{\infty} N_0 \exp\left(-\frac{nhv}{kT}\right) \\ = \int_{0}^{\infty} \frac{hv}{kT} + 2hve^{\frac{hv}{kT}} + 3hve^{\frac{3hv}{kT}} + \cdots \\ = \int_{0}^{\infty} \frac{hv}{e^{hv/kT}-1}}$$

$$Derivation \ f \ (\lambda, \tau) = \frac{2\pi\hbarc^2}{\lambda^5(e^{\hbar c/\lambda k\tau} - 1)}$$

• Energy density in the interval between v dan v+dv of the blackbody which has average energy $\langle \varepsilon \rangle = \frac{\hbar v}{e^{\hbar v/k\tau} - 1}$ can be written down in a similar manner as for the case before,

$$G(v)dv = \frac{8\pi v^2 dv}{c^3}$$

$$u(v, \tau)dv = G(v)dv \cdot \langle \varepsilon \rangle = \frac{8\pi v^2 dv}{c^3} \cdot \frac{\hbar v}{e^{\hbar v/k\tau} - 1}$$

$$hv \rightarrow \frac{hc}{\lambda}, \frac{v^2 dv}{c^3} \rightarrow \frac{d\lambda}{\lambda^4}; u(v, \tau)dv \rightarrow u(\lambda, \tau)d\lambda$$

$$u(\lambda, \tau)d\lambda = \frac{8\pi d\lambda}{\lambda^4} \cdot \frac{hc/\lambda}{e^{\hbar c/\lambda k\tau} - 1} \rightarrow R(\lambda, \tau) = \frac{c}{4}\frac{u(\lambda, \tau)d\lambda}{d\lambda} = \frac{2\pi hc^2}{\lambda^5(e^{\hbar c/\lambda k\tau} - 1)}$$

CHAPTER 3

EXPERIMENTAL EVIDENCES FOR PARTICLE-LIKE PROPERTIES OF WAVES

Photoelectricity

1

- Classically, light is treated as EM wave according to Maxwell equation
- However, in a few types of experiments, light behave in ways that is not consistent with the wave picture
- In these experiments, light behave like particle instead
- So, is light particle or wave? (recall that wave and particle are two mutually exclusive attributes of existence)
- This is a paradox that we will discuss in the rest of the course wave particle duality









$I_2 > I_1$ because more electrons are kicked out per unit time by radiation of larger intensity, *R*

- The photocurrent saturates at a larger value of I_2 when it is irradiated by higher radiation intensity R_2
- This is expected as larger *R* means energy are imparted at a higher rate on the metal surface



















Wave theory and the time delay problem

• A potassium foil is placed at a distance r = 3.5 m from a light source whose output power P_0 is 1.0 W. How long would it take for the foil to soak up enough energy (=1.8 eV) from the beam to eject an electron? Assume that the ejected electron collected the energy from a circular area of the foil whose radius is 5.3 x 10⁻¹¹ m









Wave and particle carries energy differently The way how photon carries energy is in in contrast to the way wave carries energy.

- For wave the radiant energy is continuously distributed over a region in space and not in separate bundles
- (always recall the analogy of water in a hose and a stream of ping pong ball to help visualisation)




















$E = hv_0$	Table 3.1 Photoelect Functions	Table 3.1 SomePhotoelectric WorkFunctions $W_0 = hv_0$	
333	Material	W(eV	
* * *	Na	2.28	
	Al	4.08	
	Co	3.90	
	Cu	4.70	
• KE = 0	Zn	4.31	
1	Ag	4.73	
0	Metal Pt	6.35	
	Pb	4.14	







Experimental determination of Planck constant from PE

- Experiment can measure eV_s (= K_{max}) for a given metallic surface (e.g. sodium) at different frequency of impinging radiation
- We know that the work function and the stopping potential of a given metal is given by

•
$$eV_s = hv - W_c$$





















Compton effect

• Another experiment revealing the particle nature of X-ray (radiation, with wavelength ~ 10^{-10} m)



Compton, Arthur Holly (1892-1962), American physicist and Nobel laureate whose studies of <u>X rays</u> led to his discovery in 1922 of the so-called Compton effect.

The Compton effect is the change in wavelength of high energy <u>electromagnetic</u> <u>radiation</u> when it scatters off <u>electrons</u>. The discovery of the Compton effect confirmed that electromagnetic radiation has both wave and particle properties, a central principle of <u>quantum theory</u>.

48









Photographic picture of a Compton electron

OMPTON ELECTRON

- Part of a bubble chamber picture (Fermilab'15 foot Bubble Chamber', found at the University of Birmingham). An electron was knocked out of an atom by a high energy photon.
- Photon is not shown as the photographic plate only captures the track of charged particle, not light.















PYQ 2.2 Final Exam 2003/04

Suppose that a beam of 0.2-MeV photon is scattered by the electrons in a carbon target. What is the wavelength of those photon scattered through an angle of 90°? A. 0.00620 nm B. 0.00863 nm C. 0.01106 nm D. 0.00243 nm E. Non of the above











PYQ 3(c), Final exam 2003/04

• (c) A 0.0016-nm photon scatters from a free electron. For what scattering angle of the photon do the recoiling electron and the scattered photon have the same kinetic energy?

• Serway solution manual 2, Q35, pg. 358

Solution

67

68

• The energy of the incoming photon is

 $E_{\rm i} = hc/\lambda = 0.775 {\rm ~MeV}$

- Since the outgoing photon and the electron each have half of this energy in kinetic form,
- $E_f = hc/\lambda' = 0.775 \text{ MeV} / 2 = 0.388 \text{ MeV}$ and $\lambda' = hc/E_f = 1240 \text{ eV} \cdot \text{nm} / 0.388 \text{ MeV} = 0.0032 \text{ nm}$

• The Compton shift is $\Delta \lambda = \lambda' - \lambda = (0.0032 - 0.0016) \text{ nm} = 0.0016 \text{ nm}$

- By $\Delta \lambda = \lambda_c (1 \cos \theta)$
- $= (h/m_e c) (1 \cos \theta) 0.0016 \text{ nm}$

•
$$= 0.00243 \text{ nm} (1 - \cos \theta)$$

$$\theta = 70^{\circ}$$





X-rays are simply EM radiation with very short wavelength, ~ 0.01 nm -10 nm

Some properties:

- energetic, according to $E = hc/\lambda \sim 0.1 100 \text{ keV}$ (c.f. $E \sim a$ few eV for visible light)
- travels in straight lines
- is unaffected by electric and magnetic fields
- passes readily through opaque materials highly penetrative
- causes phosphorescent substances to glow
- exposes photographic plates



PE and x-rays production happen at different energy scale

- However, both process occur at disparately different energy scale
- Roughly, for PE, it occurs at eV scale with ultraviolet radiation
- For x-ray production, the energy scale involved is much higher - at the order of 100 eV - 100 keV

73

<text><text><text>









X-ray production heats up the target material

- Due to conversion of energy from the impacting electrons to x-ray photons is not efficient, the difference between input energy, K_e and the output x-ray energy E_{γ} becomes heat
- Hence the target materials have to be made from metal that can stand heat and must have high melting point (such as Tungsten and Molybdenum)

79

Classical explanation of continuous xray spectrum:

- The continuous X-ray spectrum is explained in terms of **Bremsstrahlung:** radiation emitted when a moving electron "tekan brake"
- According to classical EM theory, an accelerating or decelerating electric charge will radiate EM radiation
- Electrons striking the target get slowed down and brought to eventual rest because of collisions with the atoms of the target material
- Within the target, many electrons collides with many atoms for many times before they are brought to rest
- Each collision causes some non-unique losses to the kinetic energy of the Bremsstrahlung electron
- As a net effect of the collective behavior by many individual collisions, the radiation emitted (a result due to the lost of KE of the electron) forms a continuous spectrum





Bremsstrahlung cannot explain

 λ_{\min}

• Notice that in the classical **Bremsstrahlung** process the x-ray radiated is continuous and there is no lower limit on the value of the wavelength emitted (because classical physics does not relate energy with wavelength). Hence, the existence of λ_{min} is not explained with the classical **Bremsstrahlung** mechanism. All range of λ from 0 to a maximum should be possible in this classical picture.

 λ_{min} can only be explained by assuming light as photons but not as EM wave ⁸³



Theoretical explanation of the experimental Value of λ_{min}

- *K* (of the Bremsstrahlung electron) is converted into the photon with $E = hc/\lambda_{min}$
- Experimentally K is caused by the external potential V that accelerates the electron before it bombards with the target, hence

$$K = eV$$

Conservation of energy requires

$$K = eV = hc/\lambda_{\min}$$

• or, $\lambda_{\min} = hc/eV = (1240 \text{ nm} \cdot \text{eV})/eV = (1240 \text{ V/V}) \text{ nm}$ which is the value measured in x-ray experiments

85

Why is λ_{\min} the same for different material?

- The production of the x-ray can be considered as an inverse process of PE
- Hence, to be more rigorous, the conservation of energy should take into account the effects due to the work potential of the target material during the emission of x-ray process, W_0
- However, so far we have ignored the effect of W_0 when we were calculating the relationship between λ_{\min} and K
- This approximation is justified because of the following reason:
- The accelerating potentials that is used to produce x-ray in a x-ray vacuum tube, V, is in the range of 10,000 V
- Whereas the work function W_0 is only of a few eV
- Hence, in comparison, W_0 is ignored wrp to eV
- This explains why λ_{\min} is the same for different target materials



















The bright spots correspond to the directions where x-rays (full ranges of wavelengths) scattered from various layers (different Braggs planes) in the crystal interfere constructively.














































PYQ 4, Test I, 2003/04

- An electron and a positron collide and undergo pair-annihilation. If each particle is moving at a speed of 0.8*c* relative to the laboratory before the collision, determine the energy of each of the resultant photon.
- **A.** 0.85MeV **B.** 1.67 MeV
- **C.** 0.51 MeV **D.** 0.72MeV
- **E.** Non of the above



































Both light and material particle display wave-particle duality

- Not only light manifest such wave-particle duality, but other microscopic material particles (e.g. electrons, atoms, muons, pions well).
- In other words:
- Light, as initially thought to be wave, turns out to have particle nature;
- Material particles, which are initially thought to be corpuscular, also turns out to have wave nature (next topic)

137









Planck constant as a measure of quantum effect

- When investigating physical systems involving its quantum nature, the theory usually involves the appearance of the constant *h*
- e.g. in Compton scattering, the Compton shift is proportional to *h*; So is photoelectricity involves *h* in its formula
- In general, when *h* appears, it means quantum effects arise
- In contrary, in classical mechanics or classical EM theory, *h* never appear as both theories do not take into account of quantum effects
- Roughly quantum effects arise in microscopic system (e.g. on the scale approximately of the order 10⁻¹⁰ m or smaller)





The postulate: there should be a symmetry between matter and wave. The wave aspect of matter is related to its particle aspect in exactly the same quantitative manner that is in the case for radiation. The total (i.e. relativistic) energy E and momentum p of an entity, for both matter and wave alike, is related to the frequency f of the wave associated with its motion via Planck constant

$$p = h/\lambda,$$

$$E = hf$$





9

Because ...





- This also means that the wave properties of matter is difficult to observe for macroscopic system (unless with the aid of some specially designed apparatus)
- The smallness of *h* in the relation $\lambda = h/p$ makes wave characteristic of particles hard to be observed
- The statement that when $h \rightarrow 0$, λ becomes vanishingly small means that:
- the wave nature will become effectively "shut-off" and appear to loss its wave nature whenever the relevant *p* of the particle is too large in comparison with the quantum scale characterised by *h*



















Detection of electron as particle destroy the interference pattern

- If in the electron interference experiment one tries to place a detector on each hole to determine through which an electron passes, the wave nature of electron in the intermediate states are destroyed
- i.e. the interference pattern on the screen shall be destroyed
- Why? It is the consistency of the wave-particle duality that demands such destruction must happen (think of the logics yourself or read up from the text)



FIGURE 4.15 Apparatus to record passage of electrons through stits. Each slit is surrounded by a loop with a meter that signals the passage of an electron through the slit. No interference lugges are seen on the screen.





Extra readings

- Those quantum enthusiasts may like to read more about wave-particle duality in Section 5.7, page 179-185, Serway, Moses and Mayer.
- An even more recommended reading on waveparticle duality: the Feynman lectures on physics, vol. III, chapter 1 (Addison-Wesley Publishing)
- It's a very interesting and highly intellectual topic to investigate











Diffraction of electron by atoms on the crystal surface

• The peak of the diffraction pattern is the $n = 1^{st}$ order

- From x-ray diffraction experiment done independently on nickle, we know *d* = 2.15 Amstrong
- Hence the wavelength corresponds to the diffraction pattern observed in the DG experiment is

$\lambda = d \sin \phi = 1.65$ Angstrom

29

• Here, 1.65 Angstrom is the experimentally inferred value, which is to be checked against the theoretical value predicted by de Broglie

Theoretical value of λ of the electron
An external potential V accelerates the electron via eV=K
In the DG experiment the kinetic energy of the electron is accelerated to K = 54 eV (non-relativistic teatment is suffice because K << m_ec² = 0.51 MeV)
According to de Broglie, the wavelength of an electron accelerated to kinetic energy of K = p²/2m_e = 54 eV (has a equivalent matter wave wavelength)
Ap = h/p = h/(2Km_e)^{1/2} = 1.67 Amstrong
In terms of the external potential, λ = h/(2eVm_e)^{1/2}
Theory's prediction matches measured value

- The result of DG measurement agrees almost perfectly with the de Broglie's prediction: 1.65 Angstrom measured by DG experiment against 1.67 Angstrom according to theoretical prediction
- Wave nature of electron is hence experimentally confirmed
- In fact, wave nature of microscopic particles are observed not only in e- but also in other particles (e.g. neutron, proton, molecules etc. – most strikingly Bose-Einstein condensate)



- Electron's de Broglie wavelength can be tuned via $\lambda = h/(2eVm_e)^{1/2}$
- Hence electron microscope can magnify specimen (x4000 times) for biological specimen or 120,000 times of wire of about 10 atoms in width





Heisenberg's uncertainty principle (Nobel Prize,1932)

- WERNER HEISENBERG (1901 1976)
- was one of the greatest physicists of the twentieth century. He is best known as a founder of <u>quantum mechanics</u>, the new physics of the atomic world, and especially for the <u>uncertainty principle</u> in quantum theory. He is also known for his controversial role as a leader of Germany's <u>nuclear fission</u> research during World War II. After the war he was active in elementary particle physics and West German <u>science policy</u>.



35

http://www.aip.org/history/heisenberg/p01. htm







Still remember the uncertainty relationships for classical waves?

• As discussed earlier, due to its nature, a wave packet must obey the uncertainty relationships for classical waves (which are derived mathematically with some approximations)

$$\Delta \lambda \Delta x > \lambda^2 \equiv \Delta k \Delta x > 2\pi \qquad \Delta t \Delta v \ge 1$$

• However a more rigorous mathematical treatment (without the approximation) gives the exact relations

$$\Delta \lambda \Delta x \ge \frac{\lambda^2}{4\pi} \equiv \Delta k \Delta x \ge 1/2 \qquad \Delta v \Delta t \ge \frac{1}{4\pi}$$

• To describe a particle with wave packet that is localised over a small region Δx requires a large range of wave number; that is, Δk is large. Conversely, a small range of wave number cannot produce a wave packet localised within a small distance.





Time-energy uncertainty

• Just as $\Delta p_x \Delta x \ge \frac{\hbar}{2}$ implies position-momentum uncertainty relation, the classical wave uncertainty relation $\Delta v \Delta t \ge \frac{1}{4\pi}$ also implies a corresponding relation between time and energy

$$\Delta E \Delta t \ge \frac{n}{2}$$

• This uncertainty relation can be easily obtained:

 $\frac{\hbar}{2}$

$$h\Delta v\Delta t \ge \frac{h}{4\pi} = \frac{\hbar}{2};$$

$$\therefore E = hv, \Delta E = h\Delta v \Longrightarrow \Delta E\Delta t = h\Delta v\Delta t =$$





- It sets the intrinsic lowest possible limits on the uncertainties in knowing the values of p_x and x, no matter how good an experiments is made
- It is impossible to specify simultaneously and with infinite precision the linear momentum and the corresponding position of a particle





- Uncertainty principle for energy.
- The energy of a system also has inherent uncertainty, ΔE
- ΔE is dependent on the *time interval* Δt during which the system remains in the given states.
- If a system is known to exist in a state of energy *E* over a limited period Δt, then this energy is uncertain by at least an amount h/(4πΔt). This corresponds to the 'spread' in energy of that state (see next page)
- Therefore, the energy of an object or system can be measured with infinite precision ($\Delta E=0$) only if the object of system exists for an infinite time ($\Delta t \rightarrow \infty$) 45



Conjugate variables (Conjugate observables)

- {*p_x*,*x*}, {*E*,*t*} are called *conjugate variables*
- The conjugate variables cannot in principle be measured (or known) to infinite precision simultaneously



Heisenberg's Gedanken experiment

- Let's say the "unperturbed" electron was initially located at a "definite" location x and with a "definite" momentum p
- When the photon 'probes' the electron it will be bounced off, associated with a changed in its momentum by some uncertain amount, Δp .
- Δp cannot be predicted but must be of the similar order of magnitude as the photon's momentum h/λ
- Hence $\Delta p \approx h/\lambda$
- The longer λ (i.e. less energetic) the smaller the uncertainty in the measurement of the electron's momentum
- In other words, electron cannot be observed without changing its momentum









Example

• A typical atomic nucleus is about 5.0×10⁻¹⁵ m in radius. Use the uncertainty principle to place a lower limit on the energy an electron must have if it is to be part of a nucleus











Example

A measurement established the position of a proton with an accuracy of ±1.00×10⁻¹¹ m. Find the uncertainty in the proton's position 1.00 s later. Assume v << c.

59

Solution • Let us call the uncertainty in the proton's position Δx_0 at the time t = 0. • The uncertainty in its momentum at t = 0is $\Delta p \geq h/(4\pi \Delta x_0)$ • Since v << c, the momentum uncertainty is $\Delta p = m \Delta v$ • The uncertainty in the proton's velocity is $\Delta v = \Delta p / m \ge h / (4\pi m \Delta x_0)$ • The distance x of the proton covers in the time t cannot be known more accurately than $\Delta x = t \Delta v \geq ht/(4\pi m \Delta x_0)$ $m=970 \text{ MeV}/c^2$ The value of Δx at t = 1.00 s is 3.15 km.





A particle contained within a finite region must has some minimal KE

- One of the most dramatic consequence of the uncertainty principle is that a particle confined in a small region of finite width cannot be exactly at rest (as already seen in the previous example)
- Why? Because...
- ...if it were, its momentum would be precisely zero, (meaning $\Delta p = 0$) which would in turn violate the uncertainty principle









CHAPTER 5 Atomic Models



Much of the luminous matter in the Universe is hydrogen. In fact hydrogen is the most abundance atom in the Universe. The colours of this Orion Nebula come from the transition between the quantized states in hydrogen atoms.

INTRODUCTION

- The purpose of this chapter is to build a simplest atomic model that will help us to understand the structure of atoms
- This is attained by referring to some basic experimental facts that have been gathered since 1900's (e.g. Rutherford scattering experiment, atomic spectral lines etc.)
- In order to build a model that well describes the atoms which are consistent with the experimental facts, we need to take into account the wave nature of electron
- This is one of the purpose we explore the wave nature of particles in previous chapters

Basic properties of atoms

- 1) Atoms are of microscopic size, $\sim 10^{-10}$ m. Visible light is not enough to resolve (see) the detail structure of an atom as its size is only of the order of 100 nm.
- 2) Most atoms are stable (i.e. atoms that are non radioactive)
- 3) Atoms contain negatively charges, electrons, but are electrically neutral. An atom with Z electrons must also contain a net positive charge of +Ze.
- 4) Atoms emit and absorb EM radiation (in other words, atoms interact with light quite readily)

Because atoms interacts with EM radiation quite strongly, it is usually used to probe the structure of an atom. The typical of such EM probe can be found in the₃ atomic spectrum as we will see now

Emission spectral lines

- Experimental fact: A single atom or molecule in a very diluted sample of gas emits radiation characteristic of the particular atom/molecule species
- The emission is due to the de-excitation of the atoms from their excited states
- e.g. if heating or passing electric current through the gas sample, the atoms get excited into higher energy states
- When a excited electron in the atom falls back to the lower energy states (de-excites), EM wave is emitted
- The spectral lines are analysed with *spectrometer*, which give important physical information of the atom/molecules by analysing the wavelengths composition and pattern of these lines.





Absorption line spectrum

 We also have absorption spectral line, in which white light is passed through a gas. The absorption line spectrum consists of a bright background crossed by dark lines that correspond to the absorbed wavelengths by the gas atom/molecules.





A successful atomic model must be able to explain the observed discrete atomic spectrum

We are going to study two attempts to built model that describes the atoms: the Thompson Plum-pudding model (which fails) and the Rutherford-Bohr model (which succeeds)

The Thompson model – Plumpudding model

Sir J. J. Thompson (1856-1940) is the Cavandish professor in Cambridge who discovered electron in cathode rays. He was awarded Nobel prize in 1906 for his research on the conduction of electricity by bases at low pressure. He is the first person to establish the particle nature of electron. Ironically his son, another renown physicist proves experimentally electron behaves like wave...



Plum-pudding model

- An atom consists of Z electrons is embedded in a cloud of positive charges that exactly neutralise that of the electrons'
- The positive cloud is heavy and comprising most of the atom's mass
- Inside a stable atom, the electrons sit at their respective equilibrium position where the attraction of the positive cloud on the electrons balances the electron's mutual repulsion



The plum-pudding model predicts unique oscillation frequency

- Radiation with frequency identical to the oscillation frequency.
- Hence light emitted from the atom in the plumpudding model is predicted to have exactly one unique frequency as given in the previous slide.
- This prediction has been falsified because observationally, light spectra from all atoms (such as the simplest atom, hydrogen,) have sets of discrete spectral lines correspond to many different frequencies (already discussed earlier).

Experimental verdict on the plum pudding model

- Theoretically one expect the deviation angle of a scattered particle by the plum-pudding atom to be small: $\Theta = \sqrt{N}\theta_{ave} \sim 1^{\circ}$
- This is a prediction of the model that can be checked experimentally
- Rutherford was the first one to carry out such experiment



FIGURE 6.2 A positively charged alpha particle is deflected by an angle θ as it passes through a Thomson-model atom. The coordinates *r* and ϕ locate the alpha particle while it is inside the atom.



FIGURE 6.6 A microscopic representation of the scattering. Some individual scatterings tend to increase Θ , while others tend to decrease Θ .

Ernest Rutherford

British physicist Ernest Rutherford, winner of the 1908 Nobel Prize in chemistry, pioneered the field of nuclear physics with his research and development of the nuclear theory of atomic structure

Born in New Zealand, teachers to many physicists who later become Nobel prize laureates

Rutherford stated that an atom consists largely of empty space, with an electrically positive nucleus in the center and electrically negative electrons orbiting the nucleus. By bombarding nitrogen gas with *alpha particles* (nuclear particles emitted through radioactivity), Rutherford engineered the transformation of an atom of nitrogen into both an atom of oxygen and an atom of hydrogen.

This experiment was an early stimulus to the development of nuclear energy, a form of energy in which nuclear transformation and disintegration release extraordinary power.



Rutherford's experimental setup

- Alpha particles from source is used to be scattered by atoms from the thin foil made of gold
- The scattered alpha particles are detected by the background screen



Copyright © 2004 Pearson Education, Inc., publishing as Addison Wesley

"...fire a 15 inch artillery shell at a tissue paper and it came back and hit you"

- In the scattering experiment Rutherford saw some electrons being bounced back at 180 degree.
- He said this is like firing "a 15-inch shell at a piece of a tissue paper and it came back and hit you"
- Hence Thompson plum-pudding model fails in the light of these experimental result

So, is the plum pudding model utterly useless?

- So the plum pudding model does not work as its predictions fail to fit the experimental data as well as other observations
- Nevertheless it's a perfectly sensible scientific theory because:
- It is a mathematical model built on sound and rigorous physical arguments
- It predicts some physical phenomenon with definiteness
- It can be verified or falsified with experiments
- It also serves as a prototype to the next model which is built on the experience gained from the failure of this model









Infrared catastrophe: insufficiency of the Rutherford model

- According to classical EM, the Rutherford model for atom (a classical model) has a fatal flaw: it predicts the collapse of the atom within 10^{-10} s
- A accelerated electron will radiate EM radiation, hence causing the orbiting electron to loss energy and consequently spiral inward and impact on the nucleus



+ 7

/ cycles/s

• The Rutherford model also cannot explain the pattern of discrete spectral lines as the radiation predicted by Rutherford model is a continuous burst.

Radiated light of

So how to fix up the problem? NEILS BOHR COMES TO THE RESCUE

- Niels Bohr (1885 to 1962) is best known for the investigations of atomic structure and also for work on radiation, which won him the 1922 Nobel Prize for physics
- He was sometimes dubbed "the God Father" in the physicist community
- http://www-gap.dcs.stand.ac.uk/~history/Mathematicians/ Bohr_Niels.html



To fix up the infrared catastrophe ...

Neils Bohr put forward a model which is a hybrid of the Rutherford model with the wave nature of electron taken into account





Postulate 2: condition for orbit stability

• Instead of the infinite orbit which could be possible in classical mechanics (c.f the orbits of satellites), it is only possible for an electron to move in an orbit that contains an integral number of de Broglie wavelengths,

•
$$n\lambda_n = 2\pi r_n, n = 1, 2, 3...$$



Electron that don't form standing

wave

- Since the electron must form standing waves in the orbits, the the orbits of the electron for each *n* is quantised
- Orbits with the perimeter that do not conform to the quantisation condition cannot persist
- All this simply means: all orbits of the electron in the atom must be quantised, and orbit that is not quantised is not allowed (hence can't exist)



Figure 4.14 A fractional number of wavelengths cannot persist because destructive interference will occur.




Third postulate

- Despite the fact that it is constantly accelerating, an electron moving in such an allowed orbit does not radiate EM energy (hence total energy remains constant)
- The atom remains stable at the ground state because there is no other states of lower energies below the ground state (hence the atom cannot make any transition to a lower energy state)
- As far as the stability of atoms is concerned, classical physics is invalid here
- My Comment: At the quantum scale (inside the atoms) some of the classical EM predictions fail (e.g. an accelerating charge radiates EM wave) 35









Important comments

- The smallest orbit charaterised by
- Z = 1, n=1 is the ground state orbit of the hydrogen

$$r_0 = \frac{4\pi\varepsilon_0 \hbar^2}{m_e e^2} = 0.5 \overset{0}{A}$$

- It's called the Bohr's radius = the typical size of an atom
- In general, the radius of an hydrogen-like ion/atom with charge Ze in the nucleus is expressed in terms of the Bohr's radius as

$$r_n = n^2 \frac{r_0}{Z}$$

- Note also that the ground state velocity of the electron in the hydrogen atom is $v_0 = 2.2 \times 10^6$ m/s << c
- non-relativistic

PYQ 7 Test II 2003/04

- In Bohr's model for hydrogen-like atoms, an electron (mass *m*) revolves in a circle around a nucleus with positive charges *Ze*. How is the electron's velocity related to the radius *r* of its orbit?
- **A.** $v = \frac{1}{4\pi\varepsilon_0} \frac{Ze^2}{mr}$ **B.** $v = \frac{1}{4\pi\varepsilon_0} \frac{Ze^2}{mr^2}$ **C.** $v = \frac{1}{4\pi\varepsilon_0} \frac{Ze}{mr^2}$
- $\mathbf{D}_{\cdot v^2} = \frac{1}{4\pi\varepsilon_0} \frac{Ze^2}{mr}$ **E.** Non of the above
- Solution: I expect you to be able to derive it from scratch without memorisation
- ANS: D, Schaum's series 3000 solved problems, Q39.13, pg 722 modified

41

Strongly recommending the Physics 2000 interactive physics webpage by the University of Colorado

For example the page

http://www.colorado.edu/physics/2000/quantu mzone/bohr.html

provides a very interesting explanation and simulation on atom and Bohr model in particular.

Please visit this page if you go online









The ground state energy

• For the hydrogen atom (Z = 1), the ground state energy (which is characterised by n = 1)

$$E_0 \equiv E_n (n=1) = -\frac{m_e e^4}{(4\pi\varepsilon_0)^2 2\hbar^2} = -13.6 \text{eV}$$

In general the energy level of a hydrogen like atom with Ze nucleus charges can be expressed in terms of

$$E_n = \frac{Z^2 E_0}{n^2} = -\frac{13.6Z^2}{n^2} \,\mathrm{eV}$$







Two important quantities to remember

- As a practical rule, it is strongly advisable to remember the two very important values
- (i) the Bohr radius, $r_0 = 0.53$ A and
- (ii) the ground state energy of the hydrogen atom, $E_0 = -13.6 \text{ eV}$

Bohr's 4th postulate explains the line spectrum



- When atoms are excited to an energy state above its ground state, they shall radiate out energy (in forms of photon) within at the time scale of $\sim 10^{-8}$ s upon their de-excitations to lower energy states –emission spectrum explained
 - When a beam of light with a range of wavelength sees an atom, the few particular wavelengths that matches the allowed energy gaps of the atom will be absorbed, leaving behind other unabsorbed wavelengthsto become the bright background in the absorption spectrum. Hence absorption spectrum explained

Example

• For example, for the H_{β} (486.1 nm) line, n = 4 in the empirical formula

$$\frac{1}{\lambda} = R_H \left(\frac{1}{2^2} - \frac{1}{n^{2_\star}} \right)$$

• According to the empirical formula the wavelength of the hydrogen beta line is

$$\frac{1}{\lambda_{\beta}} = R_{H} \left(\frac{1}{2^{2}} - \frac{1}{4^{2}} \right) = R_{H} \left(\frac{3}{16} \right) = \frac{3(1.0973732 \times 10^{7} \,\mathrm{m}^{-1})}{16}$$
$$\Rightarrow \lambda_{\beta} = 486 \,\mathrm{nm}$$

• which is consistent with the observed value















Real life example of atomic emission

 AURORA are caused by streams of fast photons and electrons from the sun that excite atoms in the upper atmosphere. The green hues of an auroral display come from oxygen



Example

- Suppose that, as a result of a collision, the electron in a hydrogen atom is raised to the second excited state (n = 3).
- What is (i) the energy and (ii) wavelength of the photon emitted if the electron makes a direct transition to the ground state?
- What are the energies and the wavelengths of the two photons emitted if, instead, the electron makes a transition to the first excited state (*n*=2) and from there a subsequent transition to the ground state?

63

Make use of $E_k = E_0 / n_k^2 = -13.6 \text{ eV} / n_k^2$ The energy of the proton emitted in the transition from the n = 3 to the n = 1 state is n = 3 $\Delta E = E_3 - E_1 = -13.6 \left(\frac{1}{3^2} - \frac{1}{1^2} \right) eV = 12.1 eV$ $\Delta E = E_3 - E_2$ the wavelength of this photon is $\lambda = \frac{c}{v} = \frac{ch}{\Delta F} = \frac{1242eV \cdot nm}{12 \ 1eV} = 102 \text{ nm}$ n = 2 $\Delta E = E_3 - E_1 \qquad \Delta E = E_2 - E_1$ Likewise the energies of the two photons emitted in the transitions from $n = 3 \rightarrow n = 2$ and $n = 2 \rightarrow n = 2$ n = 1, ground 1 are, respectively, state $\Delta E = E_3 - E_2 = -13.6 \left(\frac{1}{3^2} - \frac{1}{2^2} \right) = 1.89 \text{ eV with wavelength} \qquad \lambda = \frac{ch}{\Delta E} = \frac{1242eV \cdot nm}{1.89eV} = 657 \text{ nm}$ $\Delta E = E_2 - E_1 = -13.6 \left(\frac{1}{2^2} - \frac{1}{1^2} \right) = 10.2 \text{ eV} \text{ with wavelength } \lambda = \frac{ch}{\Delta E} = \frac{1242eV \cdot nm}{10.2eV_{64}} = 121 \text{ nm}$

Example The series limit of the Paschen $(n_f = 3)$ is 820.1 nm The series limit of a given spectral series is the shortest photon wavelength for that series The series limit of a spectral series is the wavelength corresponds to $n_i \rightarrow \infty$ What are two longest wavelengths of the Paschen series? Paschen series 1200 1400 1600 1800 2000 800 1000 Wavelength (nm) 65

Solution

- Note that the Rydberg constant is not provided
- But by definition the series limit and the Rydberg constant is closely related
- We got to make use of the series limit to solve that problem
- By referring to the definition of the series limit,

$$\frac{1}{\lambda} = R_H \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right) \xrightarrow{n_i \to \infty} \frac{1}{\lambda_\infty} = \frac{R_H}{n_f^2}$$

• Hence we can substitute $R_H = n_f^2 / \lambda_{\infty}$ into

$$\frac{1}{\lambda} = R_H \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

• and express it in terms of the series limit as $\frac{1}{\lambda} = \frac{1}{\lambda_{\infty}} \left(1 - \frac{n_f^2}{n_i^2} \right)$

•
$$n_i = 4, 5, 6...; n_f = 3$$



Example • Given the ground state energy of hydrogen atom -13.6 eV, what is the longest wavelength in the hydrogen's Balmer series? • Solution: $\Delta E = E_i - E_f = -13.6 \text{ eV} (1/n_i^2 - 1/n_f^2) = hc/\lambda$ • Balmer series: $n_f = 2$. Hence, in terms of 13.6 eV the wavelengths in Balmer series is given by $\lambda_{Balmer} = \frac{hc}{13.6 \text{eV} (\frac{1}{4} - \frac{1}{n_i^2})} = \frac{1240 \text{eV} \cdot \text{nm}}{13.6 \text{eV} (\frac{1}{4} - \frac{1}{n_i^2})} = \frac{91 \text{nm}}{(\frac{1}{4} - \frac{1}{n_i^2})}, \quad n_i = 3.45...$

$$\lambda_{Balmer} = \frac{91 \text{nm}}{\left(\frac{1}{4} - \frac{1}{n_i^2}\right)}, \ n_i = 3, 4, 5...$$

- longest wavelength corresponds to the transition from the $n_i = 3$ states to the $n_f = 2$ states
- Hence $\lambda_{Balmer,max} = \frac{91 \text{nm}}{\left(\frac{1}{4} \frac{1}{3^2}\right)} = 655.2 \text{nm}$
 - This is the red H_{α} line in the hydrogen's Balmer series
 - Can you calculate the shortest wavelength (the series limit) for the Balmer series? Ans = 364 nm



PYQ 2.18 Final Exam 2003/04

- Which of the following statements are true?
- I. the ground states are states with lowest energy
- II. ionisation energy is the energy required to raise an electron from ground state to free state
- **III**. Balmer series is the lines in the spectrum of atomic hydrogen that corresponds to the transitions to the n = 1 state from higher energy states
- A. I,IV B. I,II, IV C. I, III,IV
- **D.** I, II **E.** II,III
- ANS: D, My own question
- (note: this is an obvious typo error with the statement IV missing. In any case, only statement I, II are true.)

PYQ 1.5 KSCP 2003/04

- An electron collides with a hydrogen atom in its ground state and excites it to a state of n =3. How much energy was given to the hydrogen atom in this collision?
- **A.** -12.1 eV **B.** 12.1 eV **C.** -13.6 eV
- **D.** 13.6 eV **E.** Non of the above
- Solution: $\Delta E = E_3 - E_0 = \frac{E_0}{3^2} - E_0 = \frac{(-13.6\text{eV})}{3^2} - (-13.6\text{eV}) = 12.1\text{eV}$
- **ANS: B**, Modern Technical Physics, Beiser, Example 25.6, pg. 786

Frank-Hertz experiment The famous experiment that shows the excitation of atoms to

- The famous experiment that shows the excitation of atoms to discrete energy levels and is consistent with the results suggested by line spectra
- Mercury vapour is bombarded with electron accelerated under the potential *V* (between the grid and the filament)
- A small potential V_0 between the grid and collecting plate prevents electrons having energies less than a certain minimum from contributing to the current measured by ammeter







- As a result of inelastic collisions between the accelerated electrons of KE 4.9 eV with the Hg atom, the Hg atoms are excited to an energy level above its ground state
- At this critical point, the energy of the accelerating electron equals to that of the energy gap between the ground state and the excited state
- This is a resonance phenomena, hence current increases abruptly •
- After inelastically exciting the atom, the original (the bombarding) electron move off with too little energy to overcome the small retarding potential and reach the plate
- As the accelerating potential is raised further, the plate current again increases, since the electrons now have enough energy to reach the plate
- Eventually another sharp drop (at 9.8 V) in the current occurs because, again, the electron has collected just the same energy to excite the same energy level in the other atoms





Bohr's correspondence principle

- The predictions of the quantum theory for the behaviour of any physical system must correspond to the prediction of classical physics in the limit in which the quantum number specifying the state of the system becomes very large:
- lim quantum theory = classical theory $n \rightarrow \infty$
- At large *n* limit, the Bohr model must reduce to a "classical atom" which obeys classical theory

In other words...

• The laws of quantum physics are valid in the atomic domain; while the laws of classical physics is valid in the classical domain; where the two domains overlaps, both sets of laws must give the same result.



Example (Read it yourself)

- Classical EM predicts that an electron in a circular motion will radiate EM wave at the same frequency
- According to the correspondence principle, the Bohr model must also reproduce this result in the large *n* limit

More quantitatively

- In the limit, $n = 10^3 10^4$, the Bohr atom would have a size of 10^{-3} m
- This is a large quantum atom which is in classical domain
- The prediction for the photon emitted during transition around the $n = 10^3 - 10^4$ states should equals to that predicted by classical EM theory.













Probability of observing a photon

- Consider a beam of light
- In wave picture, $E = E_0 \sin(kx \omega t)$, electric field in radiation
- Intensity of radiation in wave picture is

$$I = \varepsilon_0 c E^2$$

- On the other hand, in the photon picture, I = Nhv
- Correspondence principle: what is explained in the wave picture has to be consistent with what is explained in the photon picture in the limit N→infinity:

$$I = \varepsilon_0 c E^2 = Nh v$$



• The probability of observing a photon at a point in unit time is proportional to *N*

• However, since $Nh\nu = \varepsilon_0 c \overline{E^2} \propto \overline{E^2}$

- the probability of observing a photon must also
- This means that the probability of observing a photon at any point in space is proportional to the square of the averaged electric field strength at that point $Prob(x) \propto E^2$

Square of the mean of the square of the wave field amplitude

<section-header>What is the physical interpretation of
matter wave?matter wave?matter wave?wave?*The wave function*, ΨIf use for the former wave?Figure 6.14 An idealized wave packet localized in space over a region Δx is the
perposition of many waves of different amplitudes and frequencies.If use wish to answer the following questions:Where is exactly the particle located within Δx ? the locality of a particle
becomes fuzzy when it's represented by its matter wave. We can no more tell
for sure where it is exactly located.The wave?Now, what does [\M 2] physically

5

Probabilistic interpretation of (the square of) matter wave

- As seen in the case of radiation field, |electric field's amplitude|² is proportional to the probability of finding a photon
- In exact analogy to the statistical interpretation of the radiation field,
- $P(x) = |\Psi|^2$ is interpreted as the probability density of observing a material particle
- More quantitatively,

mean?

Probability for a particle to be found between point a and b is

$$p(a \le x \le b) = \int_{a}^{b} P(x) dx = \int_{a}^{b} |\Psi(x,t)|^{2} dx$$





Example of expectation value: average position measured for a quantum particle

• If the position of a quantum particle is measured repeatedly with the same initial conditions, the averaged value of the position measured is given by

$$\langle x \rangle = \frac{\int_{-\infty}^{\infty} x |\Psi|^2 dx}{1} = \int_{-\infty}^{\infty} x |\Psi|^2 dx$$

Example

- A particle limited to the *x* axis has the wave function Ψ = *ax* between *x*=0 and *x*=1; Ψ = 0 else where.
- (a) Find the probability that the particle can be found between *x*=0.45 and *x*=0.55.
- (b) Find the expectation value <*x*> of the particle's position



Max Born and probabilistic interpretation

- Hence, a particle's wave function gives rise to a *probabilistic interpretation* of the position of a particle
- Max Born in 1926



German-British physicist who worked on the mathematical basis for <u>quantum mechanics</u>. Born's most important contribution was his suggestion that the <u>absolute square</u> of the wavefunction in the <u>Schrödinger equation</u> was a measure of the probability of finding the particle at a given location. Born shared the 1954 Nobel Prize in physics with <u>Bethe</u>

PYQ 2.7, Final Exam 2003/04

- A large value of the probability density of an atomic electron at a certain place and time signifies that the electron
- A. is likely to be found there
- **B.** is certain to be found there
- C. has a great deal of energy there
- D. has a great deal of charge
- E. is unlikely to be found there
- ANS:A, Modern physical technique, Beiser, MCP 25, pg. 802

Particle in in an infinite well (sometimes called particle in a box)

- Imagine that we put particle (e.g. an electron) into an "infinite well" with width *L* (e.g. a potential trap with sufficiently high barrier)
- In other words, the particle is confined within 0 < x < L
- In Newtonian view the particle is traveling along a straight line bouncing between two rigid walls









Mathematical description of standing waves In general, the equation that describes a standing wave (with a constant width *L*) is simply:

$$L = n\lambda_n/2$$

 $n = 1, 2, 3, \dots$ (positive, discrete integer)

- *n* characterises the "mode" of the standing wave
- n = 1 mode is called the 'fundamental' or the first harmonic
- n = 2 is called the second harmonics, etc.
- λ_n are the wavelengths associated with the *n*-th mode standing waves
- The lengths of λ_n is "quantised" as it can take only discrete values according to $\lambda_n = 2L/n$

Energy of the particle in the box

• Recall that

$$V(x) = \begin{cases} \infty, & x \le 0, x \ge L \\ 0, & 0 < x < L \end{cases}$$

- For such a free particle that forms standing waves in the box, it has no potential energy
- It has all of its mechanical energy in the form of kinetic energy only
- Hence, for the region 0 < x < L, we write the total energy of the particle as

 $E = K + V = p^2/2m + 0 = p^2/2m$

Energies of the particle are quantised

 Due to the quantisation of the standing wave (which comes in the form of λ_n = 2L/n), the momentum of the particle must also be quantised due to de Broglie's postulate:

$$p \rightarrow p_n = \frac{h}{\lambda_n} = \frac{nh}{2L}$$

It follows that the total energy of the particle is also quantised: $E \rightarrow E_n = \frac{p_n^2}{2m} = n^2 \frac{\pi^2 \hbar^2}{2mL^2}$

20

$$E_n = \frac{p_n^2}{2m} = n^2 \frac{h^2}{8mL^2} = n^2 \frac{\pi^2 \hbar^2}{2mL^2}$$

The n = 1 state is a characteristic state called the ground state = the state with lowest possible energy (also called zero-point energy)

$$E_n(n=1) \equiv E_0 = \frac{\pi^2 \hbar^2}{2mL^2}$$

Ground state is usually used as the reference state when we refer to ``excited states'' (n = 2, 3 or higher)

The total energy of the *n*-th state can be expressed in term of the ground state energy as

$$E_n = n^2 E_0$$
 (*n* = 1,2,3,4...)

The higher *n* the larger is the energy level


Simple analogy

• Cars moving in the right lane on the highway are in 'excited states' as they must travel faster (at least according to the traffic rules). Cars travelling in the left lane are in the ``ground state" as they can move with a relaxingly lower speed. Cars in the excited states must finally resume to the ground state (i.e. back to the left lane) when they slow down



23



 $E_{1} = (1)^{2} \frac{h^{2}}{8m_{e}L^{2}} = \frac{\left(6.63 \times 10^{-34} \,\mathrm{Js}\right)^{2}}{\left(9.1 \times 10^{-31} \,\mathrm{kg}\right)\left(100 \times 10^{-12} \,\mathrm{m}\right)^{2}} = 6.3 \times 10^{-18} \,\mathrm{J} = 37.7 \,\mathrm{eV}$







Example

A macroscopic particle's quantum state

 Consider a 1 microgram speck of dust moving back and forth between two rigid walls separated by 0.1 mm. It moves so slowly that it takes 100 s for the particle to cross this gap.
 What quantum number describes this motion?

Solution

• The energy of the particle is

$$E(=K) = \frac{1}{2}mv^{2} = \frac{1}{2}(1 \times 10^{-9} \text{ kg}) \times (1 \times 10^{-6} \text{ m/s})^{2} = 5 \times 10^{-22} \text{ J}$$

- Solving for *n* in $E_n = n^2 \frac{\pi^2 \hbar^2}{2mL^2}$
- yields $n = \frac{L}{h}\sqrt{8mE} \approx 3 \times 10^{14}$
- This is a very large number
- It is experimentally impossible to distinguish between the $n = 3 \ge 10^{14}$ and $n = 1 + (3 \ge 10^{14})$ states, so that the quantized nature of this motion would never reveal itself



PYQ 4(a) Final Exam 2003/04

- An electron is contained in a one-dimensional box of width 0.100 nm. Using the particle-in-abox model,
- (i) Calculate the *n* = 1 energy level and *n* = 4 energy level for the electron in eV.
- (ii) Find the wavelength of the photon (in nm) in making transitions that will eventually get it from the the n = 4 to n = 1 state
- Serway solution manual 2, Q33, pg. 380, modified

31

• 4a(i) In the particle-in-a-box model, standing wave is formed in the box of dimension *L*: $\lambda_n = \frac{2L}{n}$ • The energy of the particle in the box is given by $K_n = E_n = \frac{p_n^2}{2m_e} = \frac{(h/\lambda_n)^2}{2m_e} = \frac{n^2h^2}{8m_eL^2} = \frac{n^2\pi^2\hbar^2}{2m_eL^2}$ $E_1 = \frac{\pi^2\hbar^2}{2m_eL^2} = 37.7 \text{ eV}$ $E_4 = 4^2E_1 = 603 \text{ eV}$ • 4a(ii)

- The wavelength of the photon going from n = 4 to n = 1 is $\lambda = hc/(E_6 E_1)$
- = 1240 eV nm/ (603 37.7) eV = **2.2 nm**

Example on the probabilistic interpretation: Where in the well the particle spend most of its time?

- The particle spend most of its time in places where its probability to be found is largest
- Find, for the n = 1 and for n =3 quantum states respectively, the points where the electron is most likely to be found



Boundary conditions and normalisation of the wave function in the infinite well

- Due to the probabilistic interpretation of the wave function, the probability density P(x) = |Ψ|² must be such that
- $P(x) = |\Psi|^2 > 0$ for 0 < x < L
- The particle has no where to be found at the boundary as well as outside the well, i.e P(x) = |Ψ|² = 0 for x ≤ 0 and x ≥ L



$$\int_{-\infty}^{\infty} P(x) dx = \int_{0}^{L} |\Psi|^2 dx = 1$$

- is called the normalisation condition of the wave function
- It represents the physical fact that the particle is contained inside the well and the integrated possibility to find it inside the well must be 1
- The normalisation condition will be used to determine the normalisaton constant when we solve for the wave function in the Schrodinder equation

See if you could answer this question

• Can you list down the main differences between the particle-in-a-box system (infinite square well) and the Bohr's hydrogen like atom? E.g. their energies level, their quantum number, their energy gap as a function of *n*, the sign of the energies, the potential etc.

Schrodinger Equation



Schrödinger, Erwin (1887-1961), Austrian physicist and Nobel laureate. Schrödinger formulated the theory of wave mechanics, which describes the behavior of the tiny particles that make up matter in terms of waves. Schrödinger formulated the Schrödinger wave equation to describe the behavior of electrons (tiny, negatively charged particles) in atoms. For this achievement, he was awarded the 1933 Nobel Prize in physics with British physicist <u>Paul Dirac</u>

What is the general equation that governs the evolution and behaviour of the wave function?

- Consider a particle subjected to some timeindependent but space-dependent potential *V*(*x*) within some boundaries
- The behaviour of a particle subjected to a timeindependent potential is governed by the famous (1-D, time independent, non relativistic) Schrodinger equation:

$$\frac{\hbar^2}{2m}\frac{\partial^2\psi(x)}{\partial x^2} + (E - V)\psi(x) = 0$$

How to derive the T.I.S.E

- 1) Energy must be conserved: E = K + U
- 2) Must be consistent with de Brolie hypothesis that $p = h/\lambda$
- 3) Mathematically well-behaved and sensible (e.g. finite, single valued, linear so that superposition prevails, conserved in probability etc.)

41

• Read the msword notes or text books for more technical details (which we will skip here)



Infinite potential revisited

- Armed with the T.I.S.E we now revisit the particle in the infinite well
- By using appropriate boundary condition to the T.I.S.E, the solution of T.I.S.E for the wave function Ψ should reproduces the quantisation of energy level as have been deduced earlier,

i.e.
$$E_n = \frac{n^2 \pi^2 \hbar^2}{2mL^2}$$

In the next slide we will need to do some mathematics to solve for $\Psi(x)$ in the second order differential equation of TISE to recover this result. This is a more formal way compared to the previous standing waves argument which is more qualitative

Why do we need to solve the Shrodinger equation?

- The potential V(x) represents the environmental influence on the particle
- Knowledge of the solution to the T.I.S.E, i.e. $\psi(x)$ allows us to obtain essential physical information of the particle (which is subjected to the influence of the external potential V(x)), e.g the probability of its existence in certain space interval, its momentum, energies etc.

Take a classical example: A particle that are subjected to a gravity field $U(x) = GMm/r^2$ is governed by the Newton equations of motion,

$$-\frac{GMm}{r^2} = m\frac{d^2r}{dt^2}$$

• Solution of this equation of motion allows us to predict, e.g. the position of the object *m* as a function of time, r = r(t), its instantaneous momentum, energies, etc.

S.E. is the quantum equivalent of the Newton's law of motion

- The equivalent of "Newton laws of motion" for quantum particles = Shroedinger equation
- Solving for the wave function in the S.E. allows us to extract all possible physical information about the particle (energy, expectation values for position, momentum, etc.)



The behavior of the particle inside the box is governed by the equation

$$\frac{\partial^2 \psi(x)}{\partial x^2} = -B^2 \psi(x)$$

$$B^2 = \frac{2mE}{\hbar^2}$$

This term contain the information of the energies of the particle, which in terns governs the behaviour (manifested in terms of its mathematical solution) of $\psi(x)$ inside the well. Note that in a fixed quantum state *n*, *B* is a constant because *E* is conserved.

However, if the particle jumps to a state $n' \neq n, E$ takes on other values. In this case, *E* is not conserved because there is an net change in the total energy of the system due to interactions with external environment (e.g. the particle is excited by external photon)

If you still recall the elementary mathematics of second order differential equations, you will recognise that the solution to the above TISE is simply

$\psi(x) = A\sin Bx + C\cos Bx$

Where A, C are constants to be determined by ultilising the boundary conditions $_{47}$ pertaining to the infinite well system



Plug

$$\psi(x) = A \sin Bx + C \cos Bx \text{ into the LHS of EQ 2:}$$

$$\frac{\partial^2 \psi(x)}{\partial x^2} = \frac{\partial^2}{\partial x^2} [A \sin Bx + C \cos Bx]$$

$$= \frac{\partial}{\partial x} [BA \cos Bx - BC \sin Bx]$$

$$= -B^2 A \sin Bx - B^2 C \cos Bx$$

$$= -B^2 [A \sin Bx + C \cos Bx]$$

$$= -B^2 \psi(x) = \text{RHS of EQ2}$$

Proven that EQ1 is indeed the solution to EQ2

49

Boundary conditions

- Next, we would like to solve for the constants A, C in the solution ψ(x), as well as the constraint that is imposed on the constant B
- We know that the wave function forms nodes at the boundaries. Translate this boundary conditions into mathematical terms, this simply means

 $\psi(x=0)=\psi(x=L)=0$

• First,

• Plug
$$\psi(x = 0) = 0$$
 into
 $\psi = A\sin Bx + C\cos Bx$, we obtain
 $\psi(x=0) = 0 = A\sin 0 + C\cos 0 = C$

- i.e, C = 0
- Hence the solution is reduced to $\psi(x) = A \sin Bx$



- This means it must be $\sin BL = 0$, or in other words
- $B = n \pi / L = B_n, n = 1, 2, 3, ...$
- *n* is used to characterise the quantum states of $\psi_n(x)$
- B is characterised by the positive integer n, hence we use B_n instead of B
- The relationship $B_n = n\pi/L$ translates into the familiar quantisation of energy condition:

•
$$(B_n = n\pi/L)^2 \rightarrow B_n^2 = \frac{2mE_n}{\hbar^2} = \frac{n^2\pi^2}{L^2} \Longrightarrow E_n = n^2 \frac{\pi^2\hbar^2}{2mL^2}$$



Solve for
$$A_n$$
 with normalisation

$$\int_{-\infty}^{\infty} \psi_n^2(x) dx = \int_0^L \psi_n^2(x) dx = A_n^2 \int_0^L \sin^2(\frac{n\pi x}{L}) dx = \frac{A_n^2 L}{2} = 1$$
• thus
 $A_n = \sqrt{\frac{2}{L}}$
• We hence arrive at the final solution that
 $\Rightarrow \psi_n(x) = (2/L)^{1/2} \sin(n\pi x/L), n = 1, 2, 3, \dots$ for $0 < x < L$
 $\Rightarrow \psi_n(x) = 0$ elsewhere (i.e. outside the box)



Solutions

(a)
$$E_1 \equiv E_0 = \frac{\hbar^2 \pi^2}{2mL^2} = 37 \text{eV}$$
 $E_2 = n^2 E_0 = (2)^2 E_0 = 148 \text{eV}$
 $\Rightarrow \Delta E = |E_2 - E_0| = 111 \text{eV}$
(b) $P_{n=1}(x_1 \le x \le x_2) = \int_{x_1}^{x_2} \psi_0^2 dx = \frac{2}{L} \int_{x_1}^{x_2} \sin^2 \frac{\pi x}{L} dx$
For ground state $= \left(\frac{x}{L} - \frac{1}{2\pi} \sin \frac{2\pi x}{L}\right)_{x_1 = 0.09 \text{ Å}}^{x_2 = 0.11 \text{ Å}} = 0.0038$
(c) For $n = 2$, $\psi_2 = \sqrt{\frac{2}{L}} \sin \frac{2\pi x}{L}$;
 $P_{n=2}(x_1 \le x \le x_2) = \int_{x_1}^{x_2} \psi_2^2 dx = \frac{2}{L} \int_{x_1}^{x_2} \sin^2 \frac{2\pi x}{L} dx$
 $= \left(\frac{x}{L} - \frac{1}{4\pi} \sin \frac{4\pi x}{L}\right)_{x_1 = 0}^{x_2 = 0.25 \text{ Å}} = 0.25 \text{ Å}$











Why tunneling phenomena can happen

- It's due to the continuity requirement of the wave function at the boundaries when solving the T.I.S.E
- The wave function cannot just "die off" suddenly at the boundaries of a finite potential well
- The wave function can only diminishes in an exponential manner which then allow the wave function to extent slightly beyond the boundaries

$$\psi(x) = \begin{cases} A_+ \exp(Cx) \neq 0, & x \le 0\\ A_- \exp(-Cx) \neq 0, & x \ge L \end{cases}$$

- The quantum tunneling effect is a manifestation of the wave nature of particle, which is in turns governed by the T.I.S.E.
- In classical physics, particles are just particles, hence never display such tunneling effect





Figure 6.7 (a) Alpha decay of a radioactive nucleus. (b) The potential energy seen by an alpha particle emitted with energy E. R is the nuclear radius, about 10^{-14} m or 10 fm. Alpha particles tunneling through the potential barrier between R and R_1 escape the nucleus to be detected as radioactive decay products.

Real example of tunneling phenomena: alpha decay

65



Figure 3 (a) The wavefunction of an electron in the surface of the material to be studied. The wavefunction extends beyond the surface into the empty region. (b) The sharp tip of a conducting probe is brought close to the surface. The wavefunction of a surface electron penetrates into the tip, so that the electron can "tunnel" from surface to tip.

FIGURE A Highly schematic diagram of the scanning tunneling microscope process. Electrons, represented in the figure as small dots, tunnel across the gap between the atoms of the tip and sample. A feedback system that keeps the tunneling current constant causes the tip to move up and down tracing out the contours of the sample atoms.

FIGURE D An atomic force microscope scan of a stamper used to mold compact disks. The numbers given are in nm. The bumps on this metallic mold stamp out 60 nm-deep holes in tracks that are 1.6 μ m apart in the optical disks. *Photo courtesy of Digital Instruments.*